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**LEPTEN**  
LABORATÓRIOS DE ENGENHARIA DE PROCESSOS  
DE CONVERSÃO E TECNOLOGIA DE ENERGIA

# VALIDATION OF THE BRL DIFFUSE RADIATION MODEL AGAINST BSRN FLO DATA

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Boland et. al. (2010) developed a model to estimate the diffuse radiation fraction, validated for Australia. A logistic function is proposed instead of a piecewise linear or simple nonlinear functions, as made by Erbs et. al. (1982) and other authors

The present study focuses on the validation of the BRL model, as a first step for building a model for Brazilian conditions

# The BRL multiple predictor logistic model

- $$d = \frac{1}{1 + e^{(\beta_0 + \beta_1 k_t + \beta_2 AST + \beta_3 \alpha + \beta_4 K_t + \beta_5 \psi)}}$$

Persistence Factor

Daily clearness index

Solar angle (degrees)

Apparent solar time

Clearness index (minute, hourly totals)

# The BRL multiple predictor logistic model

- Multiple predictors better fit the spread of the data

- $$d = \frac{1}{1 + e^{(\beta_0 + \beta_1 k_t + \beta_2 AST + \beta_3 \alpha + \beta_4 K_t + \beta_5 \psi)}}$$

- $AST \rightarrow$  asymmetric – differences in the atmosphere between morning and afternoon
- $\alpha \rightarrow$  altitude angle (path of the sun through the atmosphere)
- $K_t \rightarrow$  the whole day may have a common characteristic
- $\psi \rightarrow \psi = \begin{cases} (k_{t-1} + k_{t+1})/2 & sr < t < ss \\ k_{t+1} & t = sunrise \\ k_{t-1} & t = sunset \end{cases}$  the inertia of the atmosphere – lagged clearness index

# The BRL multiple predictor logistic model

- BRL model  $\rightarrow$  Generic logistic model
- $\beta_0, \beta_1, \beta_2, \beta_3, \beta_4$  and  $\beta_5 \rightarrow$  calculated from amalgamating data from seven locations (Adelaide, Darwin, Maputo, Backnell, Lisbon, Macau an Uccle)

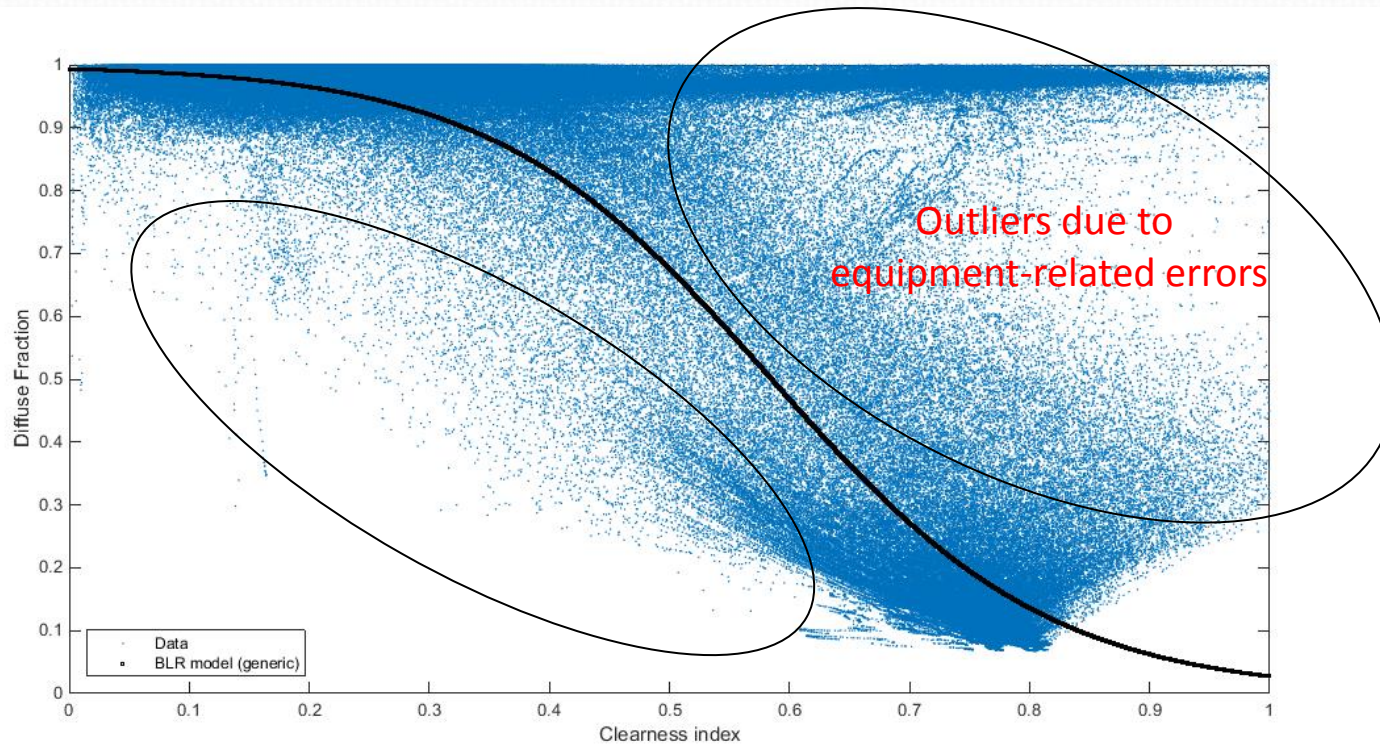
- $$d = \frac{1}{1+e^{(-5.38+6.63k_t+0.006AST-0.007\alpha+1.75K_t+1.31\psi)}}$$

# BSRN FLO

- Data from:
  - 07/1994 – 04/2002 → Qualified
  - 04/2002 – 09/2013 → Data measured, but under analysis
  - 09/2013 – 03/2016 → Qualified (renewed station)

# Results

- Considering all period
- Flagged as Qualified (Quality Control)



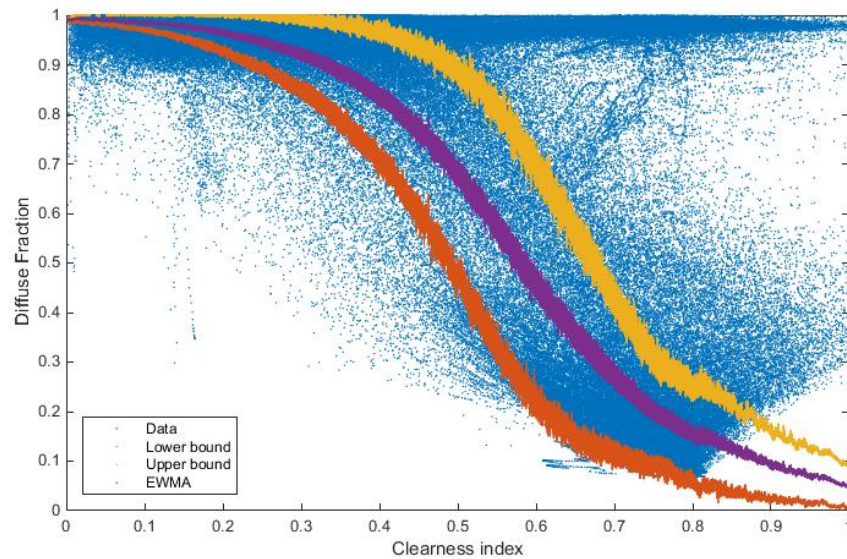
$$d = \frac{1}{1 + e^{-5.0033 + 8.6025k_t}}$$

BRL model (generic), Boland et. al. (2008,2010)

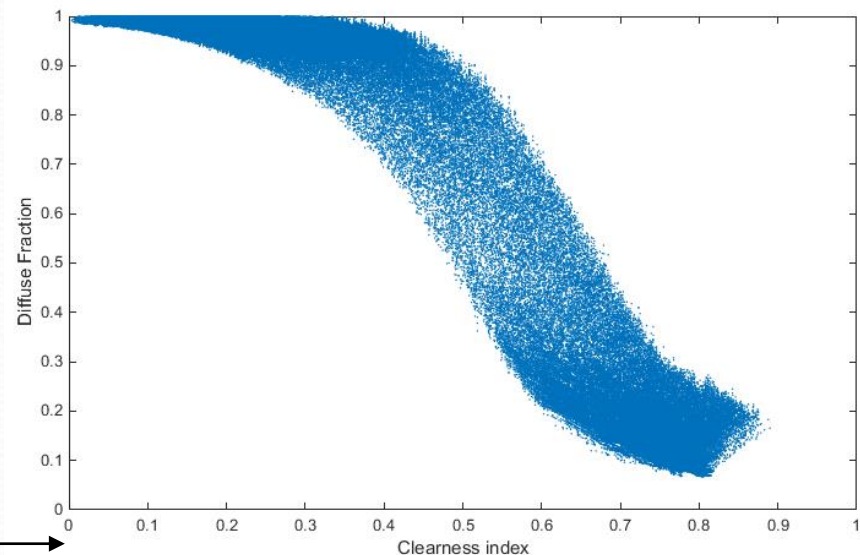
( 7 )

# Results

- Removing Outliers - “Removing values that may be wrong”
  - Quadratic programming, least squares empirical likelihood Boland et. al. (2008)
    - Unfeasible due the large amount of data
  - As suggested by Boland:
    - Used Exponentially Weighted Moving Average (EWMA) and Exponentially Weighted Moving Variance (EWMV) of the BRL model
    - Calculate the lower and upper bounds of a “Diffuse fraction vs Clearness index Envelope” Younes et. al. (2005)



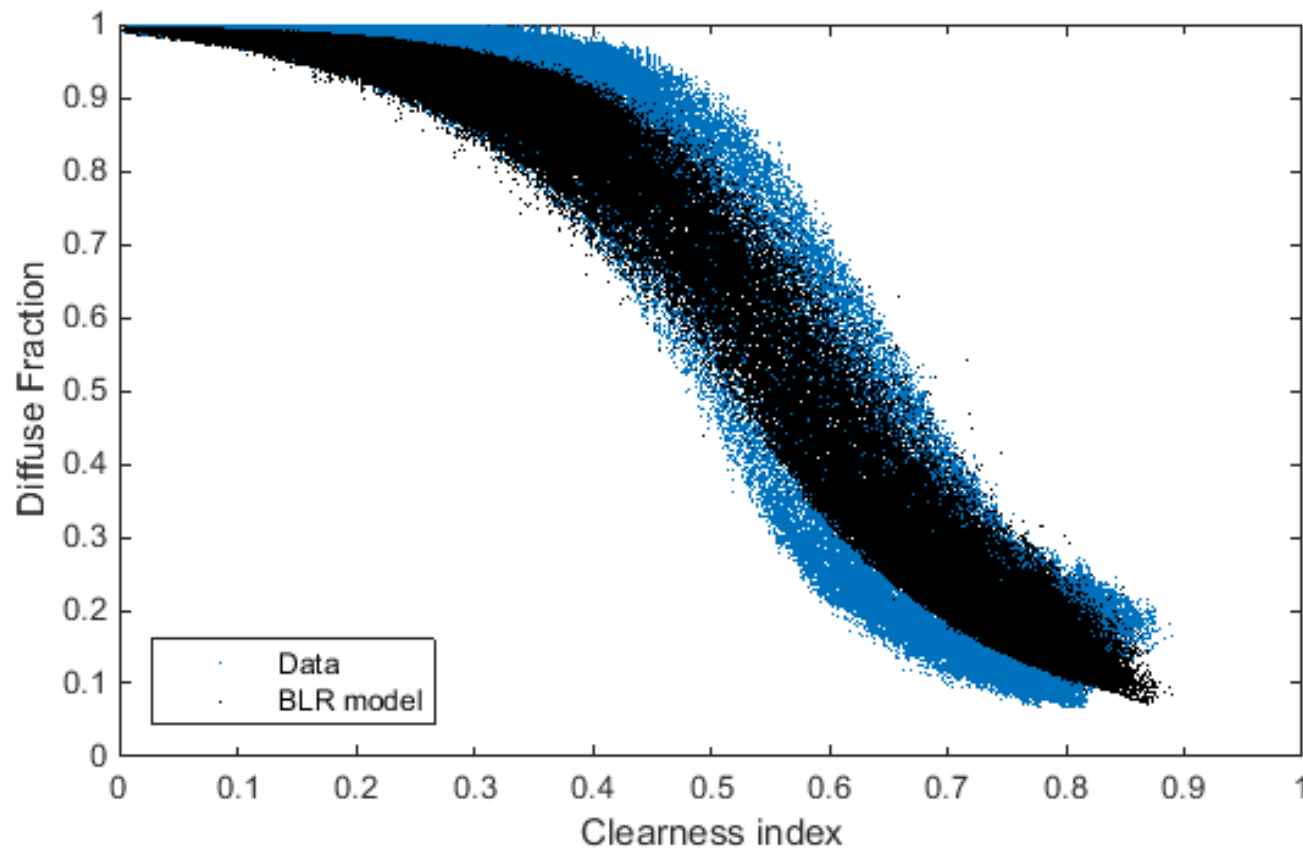
Removing the outliers





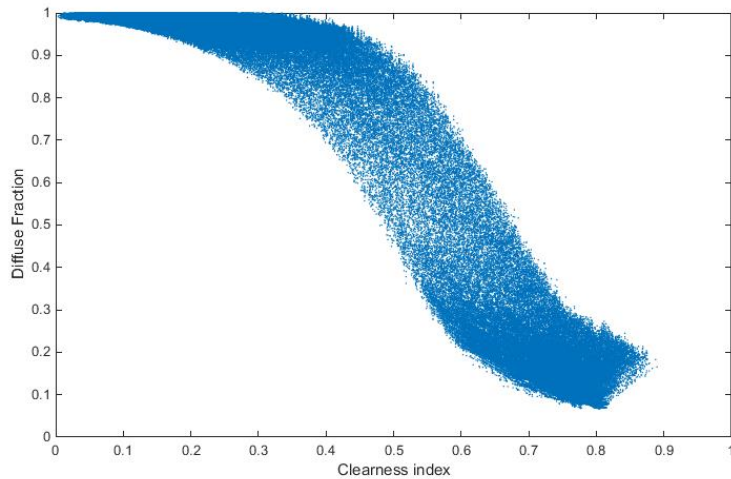
# Results

- Results using the validated BRL model

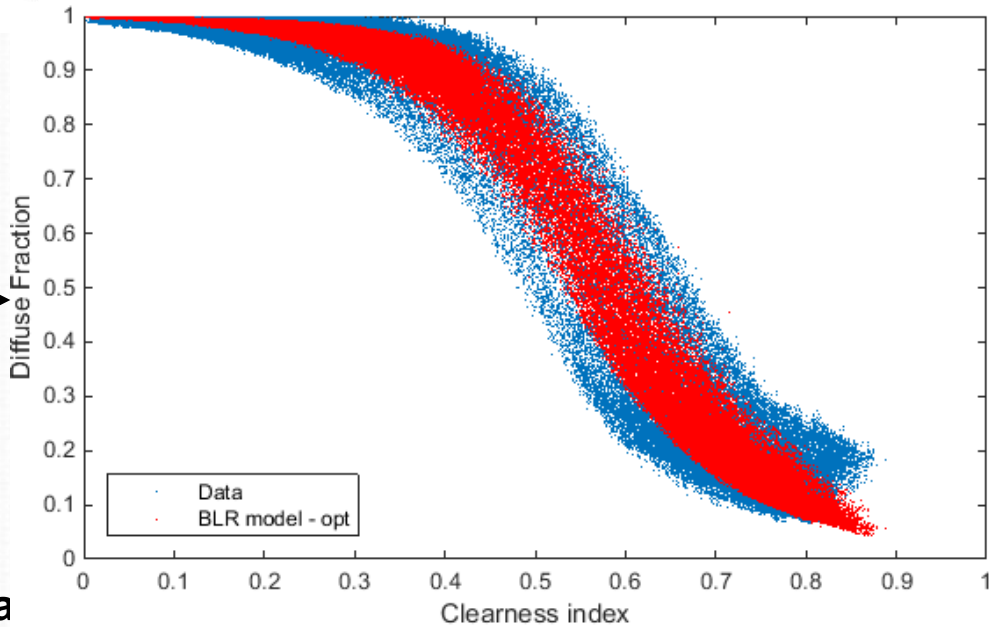
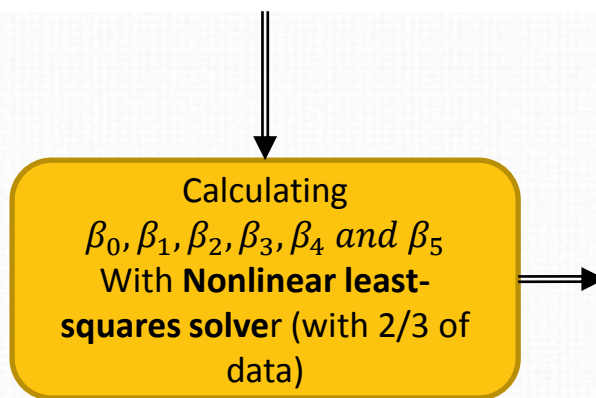


# Results

- Puling 2/3 of data to calibrate the model (random choice)

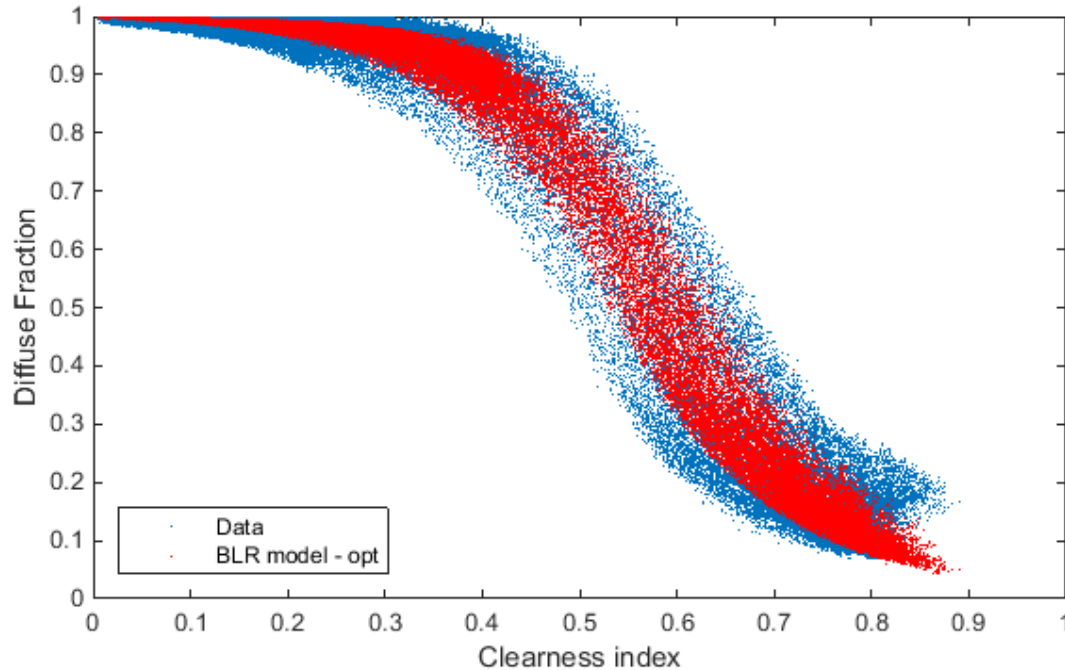


	BRL model	Model parameters for Florianopolis
$\beta_0$	-5.3796	-6.7025
$\beta_1$	6.6316	9.2319
$\beta_2$	0.006	0.0085
$\beta_3$	-0.0077	-0.0082
$\beta_4$	1.75	1.8263
$\beta_5$	1.3066	0.7875



# Results

- Validating with 1/3 of data – Error analysis



	BRL Model	BRL - FLO	
MAD (-)	0.31734	0.34343	Mean absolute deviation
MBE (-)	-0.00643	-0.0002	Mean bias error
MeAPE (%)	5.517	3.3777	Median absolute percentage error
nRMSE (%)	9.589	8.1754	Normalized root mean square error

# References

- Ridley, B., Boland, J., and Lauret, P., 2010, “Modelling of diffuse solar fraction with multiple predictors,” *Renew. Energy*, **35**(2), pp. 478–483.
- Erbs, D. G., Klein, S. A., and Duffie, J. A., 1982, “Estimation of the diffuse radiation fraction for hourly, daily and monthly-average global radiation,” *Sol. Energy*, **28**(4), pp. 293–302.
- Boland, J., and Ridley, B., 2008, “Models of diffuse solar fraction,” *Model. Sol. Radiat. Earth’s Surf. Recent Adv.*, **33**, pp. 193–219.
- Younes, S., Claywell, R., and Muneer, T., 2005, “Quality control of solar radiation data: Present status and proposed new approaches,” *Energy*, **30**(9 SPEC. ISS.), pp. 1533–1549.

THANK YOU !

